

One-parameter dependence of DeWitt's metric in Vilkoviski-DeWitt formalism

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$$\int \mathcal{D}h_\Lambda \exp \left\{ \int d^4x \mathcal{L}(\ell, h) \right\} = \exp \left\{ - \int d^4x \mathcal{L}_{eff}(\ell, \Lambda) \right\},$$

gdzie

$$\mathcal{L}(\ell, h) = \mathcal{L}_0(h) + \mathcal{L}_0(\ell) + \mathcal{L}_{int}(h, \ell) \quad \rightarrow \quad \mathcal{L}_{eff}(\ell, \Lambda) = \mathcal{L}_0(\ell) + \sum_{i \in \mathbb{N}} c_i(\Lambda) \mathcal{O}^i(\ell).$$

- Low energy dynamics (i.e., below a fundamental scale) does not depend on the details of a high energy one. The latter is encoded in coupling constants that accompany local vertices;
- Depending on the desired accuracy only a finite number of terms is required;
- Above a fundamental scale it should be sewn with a more fundamental theory;

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Gravity as an Effective Field Theory

$$\mathcal{L}(g) = \Lambda + \frac{1}{\kappa^2} R + a_1 R_{\mu\nu}^2 + a_2 R^2 + a_3 M^{-2} R^3 + \dots$$



J. F. Donoghue, *General Relativity as an effective field theory: the leading quantum corrections*, Phys.Rev.D50 (1994) 3874;

Geometry of Configuration Space

Consider the theory with a gauge symmetry generated by $K_\alpha^i(\varphi) \in G$, i.e.

$$\delta_\xi S[\varphi] = S_{,i}[\varphi] \delta_\xi \varphi^i = 0,$$

where

$$\delta_\xi \varphi^i = K_\alpha^i(\varphi) \delta \xi^\alpha, \quad K_\alpha^i(\varphi) K_{\beta,i}^j(\varphi) - K_\beta^i(\varphi) K_{\alpha,i}^j(\varphi) = f_{\alpha\beta}^\gamma(\varphi) K_\gamma^j(\varphi).$$

Let the configuration field space \mathcal{F} be a manifold endowed with a metric $\gamma(\varphi)$. Due to the underlying gauge symmetry it follows that

$$d\varphi^i = d\varphi_\perp^i + d\varphi_\parallel^i, \quad d\varphi_\parallel^i = K_\alpha^i(\varphi) d\varepsilon^\alpha \quad \text{and} \quad d\varphi_\perp^i = P_j^i d\varphi^j,$$

where P_j^i is a projector on the space orthogonal to the gauge transformations. The metric on \mathcal{F} can be cast into the form

$$ds^2 = \gamma_{ij} d\varphi^i d\varphi^j = \gamma_{ij}^\perp d\varphi_\perp^i d\varphi_\perp^j + N_{\alpha\beta} d\varepsilon^\alpha d\varepsilon^\beta,$$

where

$$N_{\alpha\beta} \equiv K_\alpha^i \gamma_{ij} K_\beta^j \quad \text{and} \quad N_{\alpha\beta} N^{\beta\gamma} = \delta_\alpha^\gamma, \quad \Rightarrow \quad P_j^i = \delta_j^i - K_\alpha^i N^{\alpha\beta} K_\beta^j \gamma_{ij},$$

The metric on the orbit space \mathcal{F}/G i.e. physical field space is the following

$$\gamma_{ij}^\perp = \gamma_{ij} - \gamma_{ik} K_\alpha^k N^{\alpha\beta} K_\beta^l \gamma_{lj}.$$

The connection that is compatible with the configuration space metric i.e. which satisfies the condition

$$\nabla_i \gamma_{jk}^\perp = 0$$

takes the form

$$\bar{\Gamma}_{kl}^i = \Gamma_{kl}^i + T_{kl}^i,$$

where Γ_{kl}^i is the connection built with the full metric γ whereas the last term is defined as

$$T_{kl}^i = -2\gamma_{(k|r} K_\alpha^r N^{\alpha\beta} D_{l)} K_\beta^s + \gamma_{(k|r} K_\alpha^r N^{\alpha\beta} K_\alpha^p D_p K_\mu^s N^{\mu\nu} K_\nu^p$$

The important property results

$$\gamma_{ij}^\perp K_\alpha^j = 0 = \gamma_{ij}^\perp \nabla_k K_\alpha^j \Rightarrow \nabla_k K_\alpha^j \sim K_\mu^j$$

The Vilkoviski-DeWitt Effective Action reads

$$e^{-\Gamma_{\text{VD}}[\bar{\phi}]} = \int \mathcal{D}\mu[\bar{\phi}; \phi] \delta[\chi] \exp \left\{ -S[\phi] - \sigma^i[\bar{\phi}, \phi] C^{-1}{}^j{}_i[\bar{\phi}] \frac{\delta \Gamma_{\text{VD}}[\bar{\phi}]}{\delta \bar{\phi}^j} \right\}$$

where $\sigma[\varphi, \phi] = 1/2(\text{geodesic connecting } \varphi \text{ and } \phi)^2$ and

$$\mathcal{D}\mu[\bar{\phi}; \phi] = \mathcal{D}\phi \sqrt{g[\bar{\phi}]} \det Q[\bar{\phi}, \phi], \quad C^{-1}{}^i{}_j[\bar{\phi}] \approx \delta^i_j + \mathcal{R}^i{}_{kjl}[\varphi] \langle \sigma^k[\bar{\phi}, \phi] \sigma^l[\bar{\phi}, \phi] \rangle + \dots$$

Quantity $\sigma^i[\bar{\phi}, \phi] \equiv g^{ij}[\bar{\phi}] \delta\sigma[\bar{\phi}, \phi] / \delta \bar{\phi}^j$ at $\bar{\phi}^i$ transforms as a vector whereas at ϕ^i as a scalar. Hence quantum gauge transformations are not affected by the presence of classical currents.

Due to the properties

$$K^k{}_\alpha[\bar{\phi}] \nabla_k \sigma^i[\bar{\phi}, \phi] \sim K^k{}_\beta[\bar{\phi}], \quad \text{and} \quad K^k{}_\alpha[\phi] \frac{\delta}{\delta \phi^k} \sigma^i[\bar{\phi}, \phi] \sim K^k{}_\beta[\bar{\phi}]$$

the VDEA is background gauge independent, *i. e.*

$$\Gamma_{\text{VD},i}[\bar{\phi}] K^i{}_\alpha[\bar{\phi}] = 0.$$

One loop approximation to VDEA

The covariant expansion of the classical action reads

$$S[\phi] = \sum_{n \geq 0} \frac{(-1)^n}{n!} (\nabla_{i_1} \dots \nabla_{i_n} S[\bar{\phi}]) \sigma^{i_1}[\bar{\phi}, \phi] \dots \sigma^{i_n}[\bar{\phi}, \phi].$$

In the one loop approximation $C^{-1}{}^i{}_j[\bar{\phi}] \approx \delta^i{}_j$. Hence,

the one loop VDEA

$$\Gamma_{VD}^{(1)}[\bar{\phi}] = \frac{1}{2} \log \det \left(\nabla_i \nabla_j S[\bar{\phi}] + \frac{1}{\alpha} \chi_{,i}^\mu[\bar{\phi}] c_{\mu\nu} \chi_{,j}^\nu[\bar{\phi}] \right) - \log \det Q[\bar{\phi}] - \frac{1}{2} \log \det \gamma(\bar{\phi}),$$

Solution to the gauge-fixing dependence: Taking another gauge-fixing term $\chi'^\alpha = \chi^\alpha + \Delta\chi^\alpha$ we get

$$\delta_\chi \Gamma_{VD}[\bar{\phi}] = -G^{ij} S_{,k} \nabla_i K_\alpha^k Q_\beta^{-1\alpha} \Delta\chi_{,j}^\beta = 0,$$

that is the effective action is **independent of the gauge fixing term**.

In "the orthogonal gauge"

$$K_\alpha^i[\bar{\phi}] g_{ij}[\bar{\phi}] \sigma^j[\bar{\phi}, \phi] = 0$$

connection simplifies to the Christoffel one *i.e.* $\bar{\Gamma}_{ij}^m = \Gamma_{ij}^m$ and $T_{kl}^i = 0$;

Gravitational Configuration Field Space

The gravitational configuration field space \mathcal{F} is endowed with one-parameter family of ultralocal metrics ($\varphi^i \rightarrow g_{\{x,\mu\nu\}} = g_{\mu\nu}(x)$)

$$\gamma_{ij}(\varphi; a) \rightarrow \sqrt{g(x)} \mathcal{G}^{\mu\nu, \alpha\beta}(x; a) \bar{\delta}(x, x') = \sqrt{g(x)} \frac{1}{4} \left(2g^{\mu(\alpha} g^{\beta)\nu} - a g^{\mu\nu} g^{\alpha\beta} \right) (x) \bar{\delta}(x, x'),$$

where $a \neq \frac{1}{2}, 2$. In the case of Einstein gravity metric can be chosen from the highest derivative term in the second order expansion of the action about a background configuration

$$S_{,ij} = \gamma_{ik} A_j^k + C_{ij} \nabla \nabla + \mathcal{B}_{ij} = P_i^k \gamma_{kl} P_j^l + \mathcal{B}_{ij},$$

where

$$\gamma_{ij} = \gamma_{ij}(\varphi; 1), \quad A_j^k \rightarrow \delta_{\mu'\nu'}^{\alpha\beta} \square \bar{\delta}(x, x'), \quad \mathcal{B}_{ij} \sim R_{..} \text{ (curvatures in background fields).}$$

Is it possible to change parametrization s.t.

$$\gamma_{ij}(\varphi; a) = \frac{\partial \varphi'^k}{\partial \varphi^i} \frac{\partial \varphi'^l}{\partial \varphi^j} \gamma_{kl}(\varphi'; 1)?$$

ANSWER: **NO**.

In principle one can choose any metric on \mathcal{F} . However, this leads to the one-parameter dependent results when VDW effective action is used. Indeed,

$$\begin{aligned} & \frac{1}{2} \log \det \left(\nabla_i(a) \nabla_j(a) S[\bar{\phi}] + \frac{1}{\alpha} \chi_{,i}^{\mu}[\bar{\phi}; a] c_{\mu\nu} \chi_{,j}^{\nu}[\bar{\phi}; a] \right) \\ &= \frac{1}{2} \log \det \left(\nabla_i \nabla_j S[\bar{\phi}] + \frac{1}{\alpha} \chi_{,i}^{\mu}[\bar{\phi}] c_{\mu\nu} \chi_{,j}^{\nu}[\bar{\phi}] \right) + (a-1) G^{ij} P_i^k H_{kl} P_j^l + \mathcal{O}((a-1)^2), \end{aligned}$$

where

$$H_{kl} \equiv 2P_k^m \gamma_{mr} Q_s^r K_{\alpha}^s N^{\alpha\beta} K_{\beta}^p D_p D_n S_j^n,$$

and

$$\gamma_{ij} = \gamma_{ik} \Pi_j^k + \gamma_{ik} Q_j^k, \quad \Pi_j^k \equiv \delta_j^k - Q_j^k, \quad Q_j^k \rightarrow \frac{1}{4} g_{\mu\nu} g^{\alpha\beta}.$$

Hence, H_{ij} does not vanish due to the presence of projector of symmetric tensor field on its trace Q_j^i .

The Scalar field interacting with Gravity

The action

$$S = -\frac{1}{\kappa^2} \int d^n x \sqrt{\bar{g}} (\bar{R} - 2\Lambda) + \int d^n x \sqrt{\bar{g}} \left(\frac{1}{2} \bar{g}^{\mu\nu} \partial_\mu \bar{\varphi} \partial_\nu \bar{\varphi} + V(\bar{\varphi}) \right),$$

where

$$V(\bar{\varphi}) = \frac{1}{2} m^2 \bar{\varphi}^2 + \frac{\lambda}{4!} \bar{\varphi}^4$$

Expansion of $S[g, \varphi]$ about the background field configuration yields:

$$\bar{g}_{\mu\nu} = g_{\mu\nu} + \kappa h_{\mu\nu} \quad \text{i} \quad \bar{\varphi} = \varphi / \kappa + \phi.$$

Hence

$$\begin{aligned} S_{,ij} \rightarrow & \frac{1}{2} h_{\mu\nu} \left(-\mathcal{G}^{\mu\nu, \alpha\beta} \square - 2\mathcal{G}^{\mu\nu, \alpha\beta} \Lambda + X_\varphi^{\mu\nu, \alpha\beta} + X_g^{\mu\nu, \alpha\beta} \right) h_{\alpha\beta} - \frac{1}{2} C_\mu^2 \\ & + \frac{1}{2} \phi \left(-\square + m^2 + \omega \phi^2 \right) \phi \quad \text{where} \quad \omega \equiv \lambda / 2\kappa^2 \\ & - h_{\mu\nu} Q^{\mu\nu, \alpha} \nabla_\alpha \phi + h_{\mu\nu} \left(\frac{1}{2} V'(\varphi) g^{\mu\nu} \right) \phi. \end{aligned}$$

Metric and Connection

The metric on \mathcal{F}

$$ds^2 = \frac{1}{\kappa^2} \int d^n x \sqrt{g} \mathcal{G}^{\mu\nu, \rho\sigma}(a) dg_{\mu\nu}(x) dg_{\rho\sigma}(x) + \int d^n x \sqrt{g} d\varphi(x) d\varphi(x)$$

The Christoffel connection on \mathcal{F}

$$\begin{aligned} \Gamma^{\mu\nu, \rho\sigma}_{\alpha\beta} &= -\delta^{\mu\nu, \rho\sigma}_{\alpha\beta} + \frac{1}{4} (g^{\rho\sigma} \delta^{\mu\nu}_{\alpha\beta} + g^{\mu\nu} \delta^{\rho\sigma}_{\alpha\beta}) + \frac{1}{2(2a-1)} g_{\alpha\beta} \mathcal{G}^{\mu\nu, \rho\sigma}(a), \\ \Gamma^{\mu\nu, 11}_{11} &= \frac{1}{4} g^{\mu\nu}, \\ \Gamma^{11, 11}_{\mu\nu} &= \kappa^2 \frac{1}{2(2a-1)} g_{\mu\nu} \end{aligned}$$

Due to a background field independence of the VDEA we take flat background metric.

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The β -function for φ^4 theory (MS)

$$\beta_\lambda(g) = \frac{1}{(4\pi)^2} \left[3\lambda^2 + 2 \left(g_\Lambda - 4g_m \frac{4a-3}{2a-1} \right) g_\kappa \lambda \right]$$

Anomalous dimension for the mass operator (MS)

$$\gamma_m = \frac{1}{(4\pi)^2} \left[\left(1 - \frac{2g_\Lambda}{g_m} \right) \lambda + \frac{g_\kappa}{2a-1} (8g_\Lambda - 5g_m) \right].$$

VD effective action and gravitational corrections

The action for the abelian YM theory

$$S = -\frac{\mu^{n-4}}{\kappa^2} \int d^n x \sqrt{g} (R - 2\Lambda) + \frac{1}{4e^2} \mu^{n-4} \int d^n x \sqrt{g} \bar{g}^{\mu\alpha} \bar{g}^{\nu\beta} \bar{F}_{\mu\nu} \bar{F}_{\alpha\beta}$$

Expanding about the background field configuration $\phi^i = \varphi^i + \eta^i$, where $\varphi^i = (g_{\mu\nu}, A_\alpha)$ and $\eta^i = (\kappa h_{\mu\nu}, a_\alpha)$ and taking one-parameter dependent metric on the full field space

$$ds^2 = \frac{1}{\kappa^2} \int d^n x \sqrt{g} \mathcal{G}^{\mu\nu, \rho\sigma}(a) dg_{\mu\nu}(x) dg_{\rho\sigma}(x) + \int d^n x \sqrt{g} g^{\alpha\beta} dA_\alpha(x) dA_\beta(x),$$

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and corresponding Christoffel connection

$$\begin{aligned} \Gamma^{\mu, \nu}_{\alpha\beta} &= \kappa^2 \delta^{\mu\nu}_{\alpha\beta}; \\ \Gamma^{\nu, \alpha\beta}_{\mu} &= -\frac{1}{2a-1} g_{\mu\lambda} \mathcal{G}^{\lambda\nu, \alpha\beta}; \end{aligned}$$

VD EA is background field independent, therefore we take $g_{\mu\nu} \rightarrow \delta_{\mu\nu}$.

The form of the β function is

$$\beta(e) = -\frac{3}{2} \frac{a}{2a-1} \frac{\Lambda \kappa^2}{(4\pi)^2} e.$$